Accelerator theory group, Accelerator laboratory, KEK

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K. Ohmi, K. Oide, Y. Funakoshi, E. Forest, D. Sagan, Y. Zhang, and SuperKEKB commissioning group

1st meeting of beam-beam workgroup, Aug. 24, 2021, KEK

Demin Zhou

# Kind reminding

- This talk gives an overview of beam-beam simulations I did in the past years in collaboration with K. Ohmi and other members of KEKB/SuperKEKB team. The main purpose is to share the lessons I learned and I hope the talk is informative to the ITF-BB workgroup as well to the ITF team.
- The "nano-beam scheme" adopted at SuperKEKB design does not include crab waist because of its strong impact on dynamic aperture and beam lifetime. Therefore, most of my old work on beam-beam simulations was done without crab waist.
- The beam-beam effects on optics design/optimization were reviewed by H. Koiso in the ITF kickoff meeting (See Ref.[1]).
- The beam-beam effects at SuperKEKB were reviewed by K. Ohmi in this meeting.







# Types of simulations for beam-beam effects [2]

- Weak-strong model + simple one-turn map: BBWS code
  - The weak beam is represented by N macro-particles (statistical errors ~  $1/\sqrt{N}$ ). The strong beam has a rigid charge distribution with its EM fields expressed by Bassetti-Erskine formula.
  - The simple one-turn map contains lattice transformation (Tunes, alpha functions, beat functions, X-Y couplings, dispersions, etc.), chromatic perturbation, synchrotron radiation damping, quantum excitation, crab waist, etc.
- Weak-strong model + full lattice: SAD code
  - The BBWS code was implemented into SAD as a type of BEAMBEAM element, where beam-beam maps is called in particle tracking.
  - Tracking using SAD: 1) Symplectic maps for elements of BEND, QUAD, MULT, CAVI, etc. 2) Element-by-element SR damping/excitation; 3) Distributed weak-strong spacecharge; 4) MAP element for arbitrary perturbation maps (such as crab waist, wake fields, artificial SR damping/excitation, etc.); ...
- Strong-strong model + simple one-turn map: BBSS code
  - Both beam are represented by N macro-particles
  - The one-turn map is the same as weak-strong code. Beamstrahlung model is also available. Choices of numerical techniques: PIC, Gaussian fitting for each slice, ...
  - For SuperKEKB, it is hard to include lattice.

 $M = M_{rad} \circ M_{chr} \circ M_{bb} \circ M_{cw} \circ M_0$ 

$$M_0 = R \cdot M_{lin} \cdot R^{-1}$$

BMBMP =(NP=3.63776D10 BEAMBEAM **BETAX=0.06 BETAY=0.001** EX=0.D0 EY=0.D0 EMIX=4.6D-9 EMIY=40.D-12 SIGZ=6.D-3 DP=6.30427D-4 ALPHAX=0.D0 ALPHAY=0.D0 DX=0.E-6 DZ=0.0 SLICE=200.D0 XANGLE=41.5D-3 **STURN=1000**)







# Types of simulations for beam-beam effects [2]

- Weak-strong model + simple one-turn map: BBWS code  $\bullet$ 
  - Pros: Fast simulation of luminosity and beam-beam effects. Not require much computing resources. Used for tune survey, fast luminosity calculation, etc..
  - Cons: Strong beam frozen. Crab waist of strong beam not implemented. Not sensitive to coherent beam-beam head-tail (BBHT) instability (BBHTI).
- Weak-strong model + full lattice: SAD code
  - Pros: Relatively fast to allow tracking with lattice. Interplay of beam-beam and lattice nonlinearity. Space-charge modeling possible. Localized geometric wakes possible.
  - Cons: Same as BBWS code. Tune survey possible but relatively slow.
- Strong-strong model + simple one-turn map: BBSS code
  - Pros: Allow dynamic evolution of 3D distribution of two beams. Detect BBHTI.
  - Cons: Tracking quite slow. Not feasible for tune survey. No effective method of parallelization.



# Beam-beam simulations with chromatic effects for KEKB

- Model of chromatic effects [3,4]
  - Twiss parameters expressed in Taylor series.
  - Chromaticities of Twiss parameters are estimated using lattice with error seeds and also measured with beams.
  - Symplectic maps for chromatic effects reconstructed and implemented into BBWS and BBSS.
- Simulations
  - BBWS: Fast survey of chromatic alpha/beta functions, and couplings. Tune survey of chromatic effects.
  - BBSS: Simulation of luminosity performance.
  - Findings: Chromatic couplings at IP causes remarkable luminosity loss at KEKB.
- Beam tunings with chromatic knobs qualitatively agreed with beam-beam simulations, contributing to luminosity boost at KEKB.
- The tools were then applied to investigate the chromatic effects on luminosity at SuperKEKB.



- BBWS simulations: Tolerances on linear and lacksquarechromatic couplings
  - Non-zero linear and chromatic couplings at IP will directly cause luminosity loss. With small emittances and beta functions at IP, luminosity of SuperKEKB will be very sensitive to these parameters.
  - Weak-strong simulations were done to define the tolerances.
  - Chromatic couplings were estimated using design lattice with error seeds.
- Findings [5]
  - Crab waist is effective in suppressing vertical blowup.
  - Crab waist is effective in increasing tolerances on linear and chromatic couplings.
  - Control of machine errors and reliable IP tuning knobs are necessary to minimize the luminosity degradation.

$$r_i(\delta) = r_{i0} + r_{i1}\delta$$
  $i = 1, 2, 3, 4$ 

Table 1: Main parameters of the SuperKEKB use for beambeam sir

simul	innulations.										
	Parameter	$e^+$	$e^-$								
	E(GeV)	4.0	7.0								
	$C(\mathbf{m})$	3016	3016								
	$N(10^{10})$	9.04	6.53								
	$eta_x^*$ (mm)	32	25								
	$\beta_{u}^{*}(mm)$	0.27	0.42								
	$\epsilon_x^{'}(nm)$	3.2	1.7								
	$\epsilon_{y}(pm)$	12.8	8.16								
	$\sigma_{z}(mm)$	6	5								
	$\sigma_{\delta}(10^{-4})$	8.34	8.34								
	$ u_x$	45.523	44.570								
	$ u_y$	43.548	41.590								
	$\tilde{\nu_z}$	0.015	0.015								

Table 2: Average and variance of the chromatic X-Y couplings calculated from 1000 seeds of errors at the SuperKEKB LER.

Parameter	Average	Varia
$r_{11}(rad)$	-0.23	0.21
$r_{21}(m)$	0.060	0.013
$r_{31}(m^{-1})$	-290	230
$r_{41}$ (rad)	37	23

### 11 0.2 xw/CW vw/CW 5x w/o CW 10.8 0.17 ິ<u>ຼິ</u> 10.6 0.14 <u><u></u><u></u><u></u><u></u></u> 0.11<sub>b</sub>> ь×10.4 0.08 10.2 0.05 w/ CW w/o CW 0.9 <mark>ې 0.8</mark> <sup>≤</sup> 0.7 0.6 0.5 -3 -2.5 -2 -1.5 -1 -0.5 0 0.5 1 1.5 2 2.5 3 r (rad)



Table 3: Tolerances for the linear and chromatic X-Y couplings at the IP of the SuperKEKB LER, assuming a rate of 20% luminosity degradation.

Parameter	w/ crab waist	w/o crab waist
$r_1^*$ (mrad)	$\pm 5.3$	$\pm 3.5$
$r_{2}^{*}\left(mm ight)$	$\pm 0.18$	$\pm 0.13$
$r_{3}^{*}(m^{-1})$	$\pm 55$	$\pm 15$
$r_4^*$ (rad)	$\pm 1.4$	$\pm 0.4$
$r_{11}$ (rad)	$\pm 2.3$	$\pm 2.0$
$r_{21}\left(m ight)$	$\pm 0.09$	$\pm 0.07$
$r_{31}(m^{-1})$	$\pm 11000$	$\pm 9400$
$r_{41}$ (rad)	$\pm 430$	$\pm 280$





- Interplay of beam-beam and lattice nonlinearity  $\bullet$ without crab waist
  - FMA was done in July of 2012 to investigate beam-beam effects on dynamic aperture [6].
  - FMA was done with ideal lattice (with DA optimization and no machine errors).
  - Particles were tracked to 1024 turns with BEAMBEAM element at IP.
- Findings [6]
  - With nominal beam-beam tune shift of ~0.09, the beam-beam interaction could strongly reduce dynamic aperture (DA).
  - Reduction of DA is more serious in LER than in HER.
  - This raised a concern of luminosity loss due to lattice nonlinearity.



# sler\_1670

sler\_1670



**Dynamic aperture (w/o beam-beam): Real space** 

 $\mathbf{X}/\sigma_{2}$ 







- Interplay of beam-beam and lattice nonlinearity
  - Simulations of luminosity were done using SAD.
  - Simulations were updated with new lattice designs when they were available.
  - KEKB was also checked.
- Findings [7]
  - Interplay of beam-beam and lattice nonlinearity might cause significant loss of luminosity at SuperKEKB (~25% loss with design lattice of LER without machine errors). This was a big surprise when reported later in the SuperKEKB ARC meeting [8].
  - The loss rate depends on lattice designs (IR model, tunes, etc.).
  - Simulations with KEKB LER lattice showed ~6% loss of luminosity. KEKB HER lattice did not show luminosity loss.

[7] D. Zhou, Crosstalk between beam-beam interaction and lattice nonlinearities in the SuperKEKB, Talk presented at the SuperKEKB optics meeting, KEK, Jan. 17, 2013. [8] D. Zhou, Beam Dynamics Issues in SuperKEKB, Talk presented at the 18th KEKB Accelerator Review Committee, KEK, Mar. 4-6, 2013.

### 8.08 6.37 σδ

εx

εγ

 $\sigma_z$ 

Design, Y. Ohnishi et al., Prog. Theor. Exp. Phys. 2012; <sup>2)</sup>sler\_1682; <sup>3)</sup>sher 5753; <sup>1)</sup>Intra-beam scattering; <sup>5)</sup>X-Y coupling, beam-beam, errors, etc.; 6) Collective effects.

simulations

HER<sup>3)</sup>

4.47

1.54

4.9

6.3

Unit

nm

pm

mm

X10-4

Bare lattice

LER<sup>2</sup>)

1.858

0.758

4.8

7.73

### Simulation results - SuperKEKB LER

SuperKEKB machine parameters

LER<sup>I</sup>)

**3.2**<sup>4)</sup>

**8.64**<sup>5)</sup>

**6**6)

HER!)

**4.6**<sup>4)</sup>

11.55)

56)

Specific luminosity: use  $\sigma_z$  of bare lattice ( $\sigma_\delta$  under-estimated  $\bullet \bullet \bullet$ 



### Simulation results - SuperKEKB HER

Specific luminosity: use σ<sub>z</sub> of design ----- ~20% loss@Design(sher\_5753) luminosity loss is mainly due to blow-up in  $\varepsilon_v$ 



### KEKB machine parameter

			Bare	Linit		
	LER	TEK'	LER <sup>2)</sup>	HER <sup>2)</sup>	Unit	
٤x	19.25	24.19	19.25	24.19	nm	
εγ	192.5	241.9	1.716	8.646	рm	
σ <sub>z</sub>	<b>4.68</b> <sup>3)</sup>	5.20 <sup>3)</sup>	4.68	5.2	mm	
σδ	7.27	6.67	7.27	6.67	×10-4	

<sup>1)</sup>Used for simulation

<sup>2)</sup>provided by Y. Ohnishi, lattice used on Jun. 17, 2009; 3) No collective effects.

~~6% loss@Beam currents on Jun.17, 2009



<sup>1</sup>Y. Funakoshi et al., Prog. Theor. Exp. Phys. 2012.





# • Findings [9]

- In addition to direct loss of luminosity, the synchro-beta resonances  $2(\nu_x - k\nu_s) = N$  become wider with lattice, limiting choice of working point.
- The chromatic terms were extracted from lattice and used in BBWS simulations. Their contribution to luminosity loss was small (this is reasonable because, with experiences of KEKB, the chromatic terms were minimized in lattice designs of SuperKEKB) compared with other unknown sources.
- Consider tune spread of particles in the physical area of - $10\sigma_x \times 10\sigma_y$ , beam-beam stretches the tune footprint in vertical direction.

Parameter	$LER(e^+)$	<b>HER</b> ( $e^-$ )
E(GeV)	4.0	7.007
$C(\mathbf{m})$	3016	3016
$N(10^{10})$	9.04	6.53
$eta_x^*$ (mm)	32	25
$\beta_{y}^{*}$ ( $\mu$ m)	0.27	0.3
$\epsilon_x$ (nm)	3.2	4.6
$\epsilon_y$ (pm)	8.64	11.5
$\sigma_z$ (mm)	6	5
$\sigma_{\delta}(10^{-4})$	8.08	6.37
$ u_x$	44.53	45.53
$ u_y$	46.57	43.57
$ u_z$	0.0247	0.028

Table 1: Main Parameters of the SuperKEKB use for **Beam-beam Simulations** 

[9] D. Zhou et al., Crosstalk Between Beam-beam Interaction and Lattice Nonlinearities in the SuperKEKB, in Proceedings of IPAC'13, Shanghai, China.







# Findings [10]

- The fast drop of specific luminosity at low bunch currents was hard to understand.
- Simple map of crab waist (perfect crab waist) was used in SAD simulations.
- Perfect crab waist is effective in recovering luminosity at high bunch currents, but not effective at low bunch currents.
- Perfect crab waist is also effective in recovering DA with beam-beam and in suppressing beam halo.
- But even perfect crab waist cannot work well in the presence of lattice nonlinearity.



FMA: LER: x-y

120

100

80<sup>ج</sup>ع

► DA w/ beam-beam

**Real space** 

-70 -50 -30 -10 10 30 50

On-momentum

Lum.: LER: BB + Crab waist ► Crab waist: simple map at IP



FMA: LER: x-y

sler\_1684

Resonance lines:

Green: 5th orde

Blue: 4th order

Tune space

0.52 0.53\_0.54 0.55 0.56 0.57 0.58







Beam tail: LER: SAD: w/ BB & CW

► LN degrade CW performance?





### sler\_1684



### sler\_1684



# • Findings [11,12]

- The power of crab waist was demonstrated in tune survey via **BBWS** simulations.
- Then space charge (SC) will play an important role (suggested) by M. Zobov). This is ONLY true if vertical emittance of  $\epsilon_v \approx 10$ pm is achieved.
- SAD simulations with weak-strong model of SC were tried and showed strong effects. The results were questionable since the used SC model was not self-consistent.
- SC effects is a topic to be investigated, requiring self-consistent simulations.



[11] D. Zhou, Beam-beam, lattice nonlinearity and space charge at SuperKEKB, Talk presented at the ICFA Mini-Workshop on Commissioning of SuperKEKB and e+e- Colliders, KEK, Nov. 11-13, 2013. [12] D. Zhou, Interplay of beam-beam, lattice nonlinearity and space charge effects in the SuperKEKB collider, in Proceedings of IPAC'15, May 3-8, 2015; Talk.



**BB** and luminosity



## **BB** and luminosity

► SuperKEKB LER w/ crab waist • Lum. tune scan by BBWS



### SC: Case of SuperKEKB LER

► Hor. tune scan



### Interplay: Baseline lattice: **BB+LN+SC**

FMA with beam distribution:  $10\sigma_x \times 10\sigma_y$ 



### Interplay: Baseline lattice: **BB+LN**









# • Findings [13,14]

- Investigations were carried out to understand the sources of luminosity in beam-beam simulations with lattice.
- H. Sugimoto made an LER lattice with IR simplified (no \_ solenoids, no offsets of QCS magnets). Luminosity loss with the simplified lattice was found to be weaker.
- It was concluded that the nonlinear IR should be the source of luminosity loss [14].
- Detuned lattices (larger  $\beta_{x,v}^*$  with respect to the final design) were prepared for the early commissioning. Beam-beam simulations using SAD did not show obvious loss of luminosity.
- It was concluded that in the early phase of SuperKEKB commissioning, lattice nonlinearity and space charge should not be show-stopper. Unfortunately, this was a misleading conclusion.

## Lum. calculation: LER: Simplified IR

- ► Simplified lattice by H. Sugimoto
- sler\_simple001.sad: no solenoid but preserve main optics parameters
- ► No significant lim. degradation at low current
- ► Solenoid is the main source of lattice nonlinearity?



L	E	R	

β <sub>v</sub> at IP	128	mm
β <sub>v</sub> at IP	2.16	mm
Ib	1	Α
Пь	2500	
εv	1.75	nm
= x E.u./E.u	2	%
cy/cx	2	-70

HER

$\beta_{X}$ at IP	100	mm
$\beta_{y}$ at IP	2.40	mm
Ι <sub>b</sub>	0.8	Α
n <sub>b</sub>	2500	
ε <sub>x</sub>	4.5	nm
ε <sub>γ</sub> /ε <sub>x</sub>	2	%

### Lum. calculation: Detuned lattice

- $\blacktriangleright$  Assume:  $\varepsilon_x$ =1.75nm, coupling = 2%
- Space-charge is not important
- ► Lattice nonlinearity is not very important
- $\blacktriangleright$  L=1 $\times$ 10<sup>34</sup>cm<sup>-2</sup>s<sup>-1</sup> is promising
- $\blacktriangleright$  L=10×10<sup>34</sup>cm<sup>-2</sup>s<sup>-1</sup> is possible by increasing beam currents?





# • Findings [12,15,16]

- To better understand the IR nonlinearity, we initiated a project of lattice translation between accelerator codes (such as PTC, Bmad, etc.) [15]. Actually this project was quite successful, creating profits to studies of other subjects and even to subjects in other accelerator projects.
- With help of E. Forest, I used PTC to calculate the nonlinear resonant driving terms (RDTs) of SuperKEKB lattice [16]. The sdependence of RDTs clearly showed the overlapping region of detector solenoids and QCS magnets generates is the source. Some 3rd RDTs are hard to be suppressed in lattice design/ optimization. For comparison, FCC-ee design (by K. Oide) had a very clean IR.

[15] D. Zhou et al., Lattice translation between accelerator simulation codes for SuperKEKB, in Proceedings of IPAC'16, Busan, Korea, May. 08-13, 2016. [16] D. Zhou, Calculation of resonance driving terms for the SuperKEKB using PTC, Talk presented at the SuperKEKB mini optics meeting, KEK, May. 26, 2016.

### 6. Future plans

> A recently initiated project: Benchmark studies for accelerator design codes

- SAD: TRISTAN, KEKB, SuperKEKB, J-PARC, ...
- Bmad: CESR, ERL, ...
- MAD/MADX: LHC, FCCs, DAφNE, Super τ-charm, ...



### **Results by PTC**

- ► PTC applied to SuperKEKB
- $2v_x v_y [(J_x)(J_y)^{1/2}]$  resonance

• This term is hard to be compensated using arc multipoles [global correction]

• This term is almost invisible in simplified and phase-1 lattices



Figure :  $|h_{20010}|$  accumulated along the ring.

## **Results by PTC**

- > PTC applied to FCC-ee t lattice: an example
- $2v_x v_y [(J_x)(J_y)^{1/2}]$  resonance for latt. ver. FCCee\_t\_65\_26
- In general, no significant 3rd resonances in FCC-ee lattices



### 6. Future plans

A good step for benchmark of SAD and Bmad • Twiss function and FMA for SuperKEKB LER



### **Results by PTC**

### ► PTC applied to SuperKEKB

- Amplitude-dependent tune shift dv<sub>x,y</sub>/dJ<sub>y,x</sub>[(J<sub>x</sub>)(J<sub>y</sub>)]
- This term is hard to be compensated using arc multipoles [global **correction**]
- This term is almost invisible in Phase-1 lattices



Figure :  $|h_{11110}|$  accumulated along the ring.

## **Results by PTC**

- > PTC applied to FCC-ee t lattice: an example
- 4th order RDTs for latt. ver. FCCee\_t\_65\_26
- Residual 4th order RDTs exist, and depend on lattice design/ optimization















# • Findings [17]

- K. Ohmi and K. Hirosawa developed a simpler method to calculated the nonlinear terms. Good agreements were found with PTC results.
- Then perturbation maps were made via MAP element in SAD to simulate luminosity loss. Finally, the term of  $p_x^2 p_y$  was found to be important. Its sources were also well understood. Other chromatic terms can also be important in addition to chromatic couplings.
- Finally we arrived at a clear picture for the luminosity loss in beam-beam simulations (weak-strong model plus design lattice): The sources are beam-beam resonances and nonlinearity of the IR. But, the remedy is far from apparent (see Ref.[1] for further information).



Figure 4: Coefficient of  $P_X^2 P_Y$  caused by skew sextupole  $(SK_2)$  and octupole  $(K_3 + SK_3)$  fields.



 $K_3 + SK_3 + Q.edge$  fields.



Figure 6: Luminosities for sextupole term (:  $P_X^2 P_Y$ ), chromatic twiss, and SAD.



- Phase-2 commissioning  $\bullet$ 
  - The Phase-2 commissioning started in March 2018 with Belle-2 detector. Lots of challenges were encountered. Beam-beam simulations were done to help understand the observed beam phenomena.
  - Observations: Peak luminosity lower than predictions of simulations; Easy blowup of one beam; Small area in tune space for good luminosity; Unexpected high Belle-2 background; No or small gain of luminosity via squeezing  $\beta_{x,y}^*$ ; Hard to approach to the design working point (.53, .57); ...
  - Tune scan using BBWS showed that the beam-beam resonances of  $\pm \nu_x + 4\nu_y + \alpha = N$  (they appear without crab waist) can be important [18].

Machine parameters of Phase-2 for beam-beam simulations

	200/6		200	0/4		0/4	100/2		
	HER	LER	HER	LER	HER	LER	HER	LER	
E (GeV)	7.007	4	7.007	4	7.007	4	7.007	4	
I₅ (mA)	285	340	285	340	285	340	285	340	
# bunch	78	39	78	39	78	39	78	39	
ε <sub>x</sub> (nm)	4.7	2.0	4.7	2.0	4.5	1.9	4.5	1.9	
ε <sub>γ</sub> (pm)	47	20	47	20	45	19	45	19	
ε <sub>z</sub> (μm)	3.7	4.5	3.7	4.5	3.4	3.5	3.4	3.6	
β <sub>x</sub> (mm)	200	200	200	200	100	100	100	100	
β <sub>y</sub> (mm)	6	6	4	4	4	4	2	2	
σ <sub>z</sub> (mm)	5.8	5.9	5.8	5.9	5.3	4.6	5.3	4.7	
Vx	45.57	44.57	45.57	44.57	45.57	44.57	45.57	44.57	
vy	43.60	46.60	43.60	46.60	43.60	46.60	43.60	46.60	
Vs	0.0234	0.0176	0.0234	0.0176	0.0258	0.0223	0.0258	0.0225	

[18] D. Zhou, Weak-strong beam-beam simulations for SuperKEKB Phase-2, Talk presented at the SuperKEKB beam-beam performance meeting, KEK, Jun. 14, 2018.

## **BBWS** simulation

➤ Optics: HER 200/4 mm and LER 200/4 mm



**Geometric luminosity:** L=4.2x1033cm-2s-1

### **Beam-beam resonances:**

 $u_x - k\nu_s = N, \quad k = 1, 2$  $2\nu_y - j\nu_s = N, \quad j = 1, 2, 3, 4$  $\nu_x + 2\nu_y + k\nu_s = N, \quad k = 1, 2, 3, 4$  $\pm \nu_x + 4\nu_y + k\nu_s = N$ 

### Lattice resonances:

$$\nu_x - \nu_y + k\nu_s = N, \quad k = -1, 0, 1$$

## **BBWS** simulation

- Optics: HER 200/4 mm and LER 200/4 mm
- Weak beam: HER: plots with normalization



**Geometric luminosity:** L=4.2x1033cm-2s-1



# Phase-2 commissioning

- In Phase-2, the working point was optimized in a triangle area between beam-beam resonances  $\pm \nu_x + 4\nu_y + \alpha = N$  and lattice resonance  $\nu_x - \nu_y + \nu_s = N$ .
- In beam commissioning, squeezing  $\beta_{x,y}^*$  did not change the situation much. The beam-beam resonances looked to be a strong constraint.
- The BBWS simulations of beam-beam resonances were consistent with D. Shatilov's simulations using Lifetrac (informed by Y. Zhang from IHEP) [18].





This plot approximately corresponds to the scheme currently adopted for SuperKEKB

[18] D. Zhou, Weak-strong beam-beam simulations for SuperKEKB Phase-2, Talk presented at the SuperKEKB beam-beam performance meeting, KEK, Jun. 14, 2018.

## **BBWS** simulation

► Optics: HER 100/2 mm and LER 100/2 mm



**Geometric luminosity:** L=7.1x10<sup>33</sup>cm<sup>-2</sup>s<sup>-1</sup>

# **BBWS** simulation

- ➤ Optics: HER 100/2 mm and LER 100/2 mm
  - Weak beam: HER: plots with normalization



**Geometric luminosity:** L=7.1x10<sup>33</sup>cm<sup>-2</sup>s<sup>-1</sup>



- Preparation of Phase-3 commissioning  $\bullet$ 
  - Before Phase-3 commissioning started, beam-beam simulations were done with "roadmap" parameter sets defined. The working point of (.57, .61) was assumed.
  - The beam-beam resonances were always of concern from viewpoint of BBWS simulations.
  - The nonzero  $\alpha$  in  $\pm \nu_x + 4\nu_y + \alpha = N$  was mysterious. It seemed to be related to  $\nu_s$  but not confirmed.

	1		1ex		2	2		ex	3		3'		
	HER	LER	HER	LER	HER	LER	HER	LER	HER	LER	HER	LER	HE
I <sub>b</sub> (A)	1.0	1.2	1.0	1.4	1.0	1.4	1.2	1.7	1.3	1.8	1.15	1.6	1.
# bunch	15	76	15	76	15	76	15	76	15	76	15	76	
ε <sub>x</sub> (nm)	4.6	2.0	4.6	2.0	4.6	2.0	4.6	2.0	4.6	2.0	4.6	2.0	4.
ε <sub>y</sub> (pm)	368	160	230	150	138	140	128.8	130	138	140	101.2	100	101
β <sub>x</sub> (mm)	100	100	100	100	100	100	100	100	100	100	100	100	10
β <sub>y</sub> (mm)	3	3	3	3	2	2	2	2	1.4	1.4	1.25	1.25	1.3
σ <sub>z</sub> (mm)	6	6	6	6	6	6	6	6	6	6	6	6	6
٧x	45.57	44.57	45.57	44.57	45.57	44.57	45.57	44.57	45.57	44.57	45.57	44.57	45.
Vy	43.61	46.61	43.61	46.61	43.61	46.61	43.61	46.61	43.61	46.61	43.61	46.61	43.
٧s	0.0258	0.0225	0.0258	0.0225	0.0258	0.0225	0.0258	0.0225	0.0258	0.0225	0.0258	0.0225	0.02
ξ <sub>y</sub> (Geom.)	0.0272	0.0262	0.0328	0.0331	0.0278	0.0351	0.0351	0.0436	0.0302	0.0387	0.0301	0.0397	0.03
£(Geom.)	1.06	E+34	I.46	E+34	2.08	2.08E+34		3.14E+34		4.11E+34		4.00E+34	
£(BBSS)	1.00	E+34	1.30	E+34	1.74	E+34	2.16	E+34	2.52	E+34	2.55	E+34	3

"Roadmap" parameter sets for Phase-3

## **BBWS simulation: Tune scan**



## **BBWS simulation: Tune scan**





e+(S)e-(W)





- Preparation of Phase-3 commissioning
  - Meanwhile I started to run strong-beam beam-beam simulations using BBSS. Turned out that for all the "roadmap" parameter sets, the coherent BBHTI (it was discovered by K. Ohmi in 2016) [20,21]) occurs [19].
  - The BBHTI is hard to appear in BBWS simulations. Strong-strong simulations are necessary to detect BBHTI. But BBSS simulations are time-consuming and require lots of CPUs.
  - The BBHTI was believed to be avoidable provided that working point was properly chosen. BBHTI looked to be invisible or weak in beam commissioning, supporting this belief.

	1		1ex			2		2ex		3		3'		3ex	
	HER	LER	HER	LER	HER	LER	HER	LER	HER	LER	HER	LER	HER	LER	
I <sub>b</sub> (A)	1.0	1.2	1.0	1.4	1.0	1.4	1.2	1.7	1.3	1.8	1.15	1.6	1.4	2.0	
# bunch	15	76	15	76	15	76	15	76	15	76	15	76	15	76	
ε <sub>x</sub> (nm)	4.6	2.0	4.6	2.0	4.6	2.0	4.6	2.0	4.6	2.0	4.6	2.0	4.6	2.0	
ε <sub>y</sub> (pm)	368	160	230	150	138	140	128.8	130	138	140	101.2	100	101.2	100	
β <sub>x</sub> (mm)	100	100	100	100	100	100	100	100	100	100	100	100	100	100	
β <sub>y</sub> (mm)	3	3	3	3	2	2	2	2	1.4	1.4	1.25	1.25	1.2	1.2	
σ <sub>z</sub> (mm)	6	6	6	6	6	6	6	6	6	6	6	6	6	6	
٧x	45.57	44.57	45.57	44.57	45.57	44.57	45.57	44.57	45.57	44.57	45.57	44.57	45.57	44.57	
Vy	43.61	46.61	43.61	46.61	43.61	46.61	43.61	46.61	43.61	46.61	43.61	46.61	43.61	46.61	
Vs	0.0258	0.0225	0.0258	0.0225	0.0258	0.0225	0.0258	0.0225	0.0258	0.0225	0.0258	0.0225	0.0258	0.022	
ξ <sub>y</sub> (Geom.)	0.0272	0.0262	0.0328	0.0331	0.0278	0.0351	0.0351	0.0436	0.0302	0.0387	0.0301	0.0397	0.0369	0.045	
£(Geom.)	1.06	E+34	1.46	E+34	2.08	E+34	3.14E+34		4.11E+34		4.00E+34		6.20E+34		
L(BBSS)	1.00	E+34	1.30	E+34	1.74	E+34	2.16	E+34	2.52	E+34	2.55	E+34	3.21	E+34	

"Roadmap" parameter sets fo Phase-3 [22]

[19] D. Zhou, Beam-beam simulations for SuperKEKB Phase-3, Talk presented at the SuperKEKB optics meeting, KEK, Dec. 13, 2018. [20] K. Ohmi et al., Coherent Beam-Beam Instability in Collisions with a Large Crossing Angle, Phys. Rev. Lett. 119, 134801 (2017). [21] N. Kuroo et al., Cross-wake force and correlated head-tail instability in beam-beam collisions with a large crossing angle, Phys. Rev. Accel. Beams 21, 031002 (2018).



### e+ beam



### e+ beam

Param. set (1

Param. set (2)

Param. set (1ex)

Param. set (2ex) Param. set (3)

Param. set (3

Param, set (3ex

1500

1400

1300

1200

1100

Ê1000

b<sup>≻</sup>900

800

700

600 500

400



e- beam







- Preparation of Phase-3 commissioning  $\bullet$ 
  - To avoid BBHTI, it is preferred to have same fractional  $\nu_{x,y,s}$  for the two colliding beams [22]. But it is hard to be satisfied in realistic beam commissioning.
  - The BBHTI is serious around synchro-beta resonances of  $\nu_x - k\nu_s = N/2$ , therefore unequal  $\nu_s$  of the two beams will set narrower choices of  $\nu_x$  [22].
  - The BBWS and BBSS simulations have consistency when BBHTI is not important.

	-		10	ex	2	2	20	ex	:	3	3'		
	HER	LER	HER	LER	HER	LER	HER	LER	HER	LER	HER	LER	HE
I₀ (A)	1.0	1.2	1.0	1.4	1.0	1.4	1.2	1.7	1.3	1.8	1.15	1.6	1.
# bunch	15	76	15	76	15	76	15	76	15	76	1576		
ε <sub>x</sub> (nm)	4.6	2.0	4.6	2.0	4.6	2.0	4.6	2.0	4.6	2.0	4.6	2.0	4.
ε <sub>y</sub> (pm)	368	160	230	150	138	140	128.8	130	138	140	101.2	100	10
β <sub>x</sub> (mm)	100	100	100	100	100	100	100	100	100	100	100	100	10
β <sub>y</sub> (mm)	3	3	3	3	2	2	2	2	1.4	1.4	1.25	1.25	1.
σ <sub>z</sub> (mm)	6	6	6	6	6	6	6	6	6	6	6	6	6
Vx	45.57	44.57	45.57	44.57	45.57	44.57	45.57	44.57	45.57	44.57	45.57	44.57	45.
Vy	43.61	46.61	43.61	46.61	43.61	46.61	43.61	46.61	43.61	46.61	43.61	46.61	43.
Vs	0.0258	0.0225	0.0258	0.0225	0.0258	0.0225	0.0258	0.0225	0.0258	0.0225	0.0258	0.0225	0.02
ξ <sub>y</sub> (Geom.)	0.0272	0.0262	0.0328	0.0331	0.0278	0.0351	0.0351	0.0436	0.0302	0.0387	0.0301	0.0397	0.03
£(Geom.)	1.06	E+34	I.46	E+34	2.08	E+34	3.14	E+34	4.11E+34		4.00E+34		6
£(BBSS)	1.00	E+34	1.30	E+34	1.74	E+34	2.16	E+34	2.52	E+34	2.55	E+34	3

"Roadmap" parameter sets for Phase-3 [22]



## **BBSS** simulation

- > All parameter set (3ex):  $v_y = *.61$ 
  - Scan of v<sub>x</sub> (same fractional part for LER and HER)





## Beam sizes for v<sub>s+</sub>=.0225, v<sub>s-</sub>=.0258

## **BBSS** simulation

## All parameter set (3ex): vy=\*.61

• Synchro-beta resonances are wider in BBSS simulation than in those in **BBWS** 

• The luminosity slope (black arrow) can be explained in BBWS sim.



- Preparation of Phase-3 commissioning ullet
  - The beam-beam resonances seen in BBWS simulations also appear in BBSS simulations. The main difference is from BBHTI [23].

			1(op1)		1(op2)		2019.03.30		2019.04.02		2018.07.13(1	
	HER	LER	HER	LER	HER	LER	HER	LER	HER	LER	HER	L
I₅ (A)	1.0	1.2	1.0	1.2	1.0	1.2	0.21	0.26	0.17	0.22	0.158	-
# bunch	15	76	15	76	15	76	78	39	78	89	39	95
ε <sub>x</sub> (nm)	4.6	2.0	4.6	2.0	4.6	2.0	4.728	1.731	4.537	1.641	4.6	1
ε <sub>y</sub> (pm)	368	160	160	160	160	160	122.5	40	53.33	13.33	42.4	4
β <sub>x</sub> (mm)	100	100	100	230	80	80	200	200	100	200	100	2
β <sub>y</sub> (mm)	3	3	3	3	3	3	4	4	3	3	3	
σ <sub>z</sub> (mm)	6	6	6	6	6	6	6	6	6	6	6	
σ <sub>y</sub> (nm)	1051	693	693	693	693	693	700	400	400	200	357	3
Vx	45.57	44.57	45.57	44.57	45.57	44.57	45.564	44.571	45.5439	44.5568	45.542	44
Vy	43.61	46.61	43.61	46.61	43.61	46.61	43.603	46.610	43.6082	46.618	43.6072	46
Vs	0.0258	0.0225	0.0258	0.0225	0.0258	0.0225	0.0256	0.0219	0.02576	0.02205	0.0258	0.0
ξ <sub>y</sub> (Geom.)	0.0272	0.0262	0.0272	0.0397	0.0272	0.0397	0.0272	0.0220	0.0345	0.0233	0.0309	0.0
£(Geom.)	1.06	E+34	1.36	I.36E+34		E+34	1.50E+33		I.85E+33		2.46E+3	
£(BBSS)	1.00	E+34	9.30	E+33	1.34	E+34	1.25	E+33	1.39	E+33	2.43	E+3
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## Phase-2 (Last stage) and Phase-3 machine parameters (Early stage)

"Roadmap" and operation parameter sets [23]

## **BBSS** simulation

> Parameter set (1):  $v_y = *.61$ 

• Scan of v<sub>x</sub> (same fractional part for LER and HER)

### BBSS BBWS (e+ weak) 📥 BBWS (e- weak) — <u>, 0.6</u> 0.2 0.5 0.56 Fractional v 0.6 0.62



## **BBSS** simulation



• Scan of v<sub>y</sub> (same fractional part for LER and HER)









- Phase-3 commissioning  $\bullet$ 
  - The beam-beam resonances seen in BBWS simulations also appear in BBSS simulations. The main difference is from BBHTI [23].
  - I also started to take parameter sets observed in the control room for simulations.
  - It was found that in simulations unequal  $\nu_{\gamma}$  of the two beams is worse than equal  $\nu_{\chi}$  [24].

	2019.03.30		2019.04.02		2019.06.20*		2019.07.01	
	HER	LER	HER	LER	HER	LER	HER	LER
I₀ (A)	0.21	0.26	0.17	0.22	0.5	0.5	0.8	0.8
# bunch	7	89	78	89	15	76	15	76
ε <sub>x</sub> (nm)	4.728	1.731	4.537	1.641	4.537	1.64	4.49	1.93
ε <sub>γ</sub> (pm)**	122.5	40	53.33	13.33	7.5	18	16.2	6.05
β <sub>x</sub> (mm)	200	200	100	200	100	200	80	80
β <sub>y</sub> (mm)	4	4	3	3	3	3	2	2
σ <sub>z</sub> (mm)	6	6	6	6	5.5	5	5.5	5.2
σ <sub>y</sub> (nm)	700	400	400	200	150	230	180	110
Vx	45.564	44.571	45.5439	44.5568	45.5439	44.559	45.53	44.542
Vy	43.603	46.610	43.6082	46.618	43.6082	46.618	43.583	46.605
Vs	0.0256	0.0219	0.02576	0.02205	0.02718	0.023	0.02717	0.0234
y(Geom.)	0.0272	0.0220	0.0345	0.0233	0.041	0.1	0.088	0.089
C(Geom.)	1.50	E+33	1.85E+33		I.16E+34		3.78E+34	

## > Phase-3 (Early stage) machine parameters

• A few examples of parameter sets observed in the control room

**Operation parameter** sets [24]

[23] D. Zhou, Recent results of beam-beam simulations for SuperKEKB, Talk presented at the 2nd SuperKEKB beam dynamics workgroup meeting, KEK, Apr. 16, 2019. [24] D. Zhou, Simulations of beam-beam effects, Talk presented at the 1st SuperKEKB Beam Dynamics Mini-Workshop, KEK, Jul. 17, 2019.

## **BBSS** simulation

- > Parameter set (1):  $v_x = *.56$ 
  - Scan of vy (same fractional part for LER and HER)
  - Beam very unstable for v<sub>v</sub><\*.53



## **BBSS** simulation

- Parameter set (2019.04.02)
  - Scan of v<sub>x</sub>: Comparison of strong-strong and weak-strong simulations)
  - Away from resonances, BBWS works well.





3BSS (e+ beam)



### Phase-3 commissioning $\bullet$

- The tolerances of luminosity and beam-size blowup on closed orbit and linear IP aberrations were also simulated using BBWS [24].
- It was found that small vertical emittances will make the \_ luminosity and vertical beam sizes to be very sensitive to vertical orbit offset and vertical crossing angle [24].

2019.03.30		2019.04.02		2019.	06.20*	2019.07.01	
HER	LER	HER	LER	HER	LER	HER	LER
0.21	0.26	0.17	0.22	0.5	0.5	0.8	0.8
71	89	789		15	76	15	76
4.728	1.731	4.537	1.641	4.537	1.64	4.49	1.93
122.5	40	53.33	13.33	7.5	18	16.2	6.05
200	200	100	200	100	200	80	80
4	4	3	3	3	3	2	2
6	6	6	6	5.5	5	5.5	5.2
700	400	400	200	150	230	180	110
45.564	44.571	45.5439	44.5568	45.5439	44.559	45.53	44.542
43.603	46.610	43.6082	46.618	43.6082	46.618	43.583	46.605
0.0256	0.0219	0.02576	0.02205	0.02718	0.023	0.02717	0.02349
0.0272	0.0220	0.0345	0.0233	0.041	0.1	0.088	0.089
1.50	E+33	I.85E+33		I.16E+34		3.78E+34	
	2019. HER 0.21 72 4.728 122.5 200 4 6 700 45.564 43.603 0.0256 0.0272 1.50	2019.03.30       HER     LER       0.21     0.26       789       4.728     1.731       122.5     40       200     200       4     4       6     6       700     400       45.564     44.571       43.603     46.610       0.0256     0.0219       0.0272     0.0220	2019.03.30 $2019.$ HERLERHER $0.21$ $0.26$ $0.17$ $78$ $0.26$ $0.17$ $78$ $1.731$ $4.537$ $4.728$ $1.731$ $4.537$ $122.5$ $40$ $53.33$ $200$ $200$ $100$ $4$ $4$ $3$ $6$ $6$ $6$ $700$ $400$ $400$ $45.564$ $44.571$ $45.5439$ $43.603$ $46.610$ $43.6082$ $0.0256$ $0.0219$ $0.02576$ $0.0272$ $0.0220$ $0.0345$ $1.50E+33$ $1.851$	2019.03.302019.04.02HERLERHERLER0.210.260.170.22 $789$ 7897894.7281.7314.5371.641122.54053.3313.33200200100200443366670040040020045.56444.57145.543944.556843.60346.61043.608246.6180.02560.02190.025760.022050.02720.02200.03450.02331.50±+331.85±+331.85±+33	2019.04.022019.04.022019.04.022019.04.02HERLERHERLERHER0.210.260.170.220.5 $78$ 78781.6414.5374.7281.7314.5371.6414.537122.54053.3313.337.5200200100200100443336665.570040040020015045.56444.57145.543944.556845.543943.60346.61043.608246.61843.60820.02560.02190.025760.022050.027180.02720.02200.03450.02330.041	2019.03.302019.04.022019.06.20*HERLERHERLERHERLER0.210.260.170.220.50.5 $78$ $78$ $1.541$ 4.5371.6414.5371.64122.54053.3313.337.5182002001002001002004433336665.5570040040020015023045.56444.57145.543944.556845.543944.55943.60346.61043.608246.61843.608246.6180.02560.02190.025760.022050.027180.0230.02720.02200.03450.02330.0410.11.50 \= 331.85 \= +331.16 \= +341.16 \= +34	2019.03.302019.04.022019.06.20°2019.0HERLERHERLERHERLERHER0.210.260.170.220.50.50.8 $789$ $789$ $1576$ 154.7281.7314.5371.6414.5371.644.49122.54053.3313.337.51816.22002001002001002008044333266665.555.570040040020015023018045.56444.57145.543944.556845.543944.55945.5343.60346.61043.608246.61843.608246.61843.5830.02560.02190.025760.02330.0410.10.0881.50 $\pm 33$ 1.85 $\pm 33$ 1.85 $\pm 33$ 1.16 $\pm 34$ 3.78

## Phase-3 (Early stage) machine parameters

A few examples of parameter sets observed in the control room

**Operation parameter** sets [24]

[24] D. Zhou, Simulations of beam-beam effects, Talk presented at the 1st SuperKEKB Beam Dynamics Mini-Workshop, KEK, Jul. 17, 2019.

## **Tolerances of IP aberrations**

## ► Various IP aberrations: the case of e+ beam

• DPY\* (vert. crossing angle)

\*\* Luminosity becomes very sensitive to DPY\* at very small vertical emittance. This is due to increase of projected vertical emittance

\*\* Luminosity drops faster around DPY\*=0 because of additional blowup due to nonzero v-angle

\*\* With strong beam-beam blowup, the dependence can be "Tri Peaks"



## **Tolerances of IP aberrations**

### Various IP aberrations: the case of e+ beam

- DPY\* (vert. crossing angle)
- **\*\*** With small blowup at DPY\*=0, the dependence is "V" shape
- \*\* With strong blowup at DPY\*=0, the dependence becomes "M" shape
- \*\* With strong blowup at DPY\*=0, the vertical beam size is sensitive to DPY\* around 0





### Phase-3 commissioning $\bullet$

- The tolerances of luminosity and beam-size blowup on closed orbit and linear IP aberrations were also simulated using BBWS [24].
- It was found that small vertical emittances will make the \_ luminosity and vertical beam sizes to be very sensitive to vertical orbit offset and vertical crossing angle [24].

LIED						2019.07.01	
	LER	HER	LER	HER	LER	HER	LER
0.21	0.26	0.17	0.22	0.5	0.5	0.8	0.8
78	39	78	789		76	15	76
4.728	1.731	4.537	1.641	4.537	1.64	4.49	1.93
122.5	40	53.33	13.33	7.5	18	16.2	6.05
200	200	100	200	100	200	80	80
4	4	3	3	3	3	2	2
6	6	6	6	5.5	5	5.5	5.2
700	400	400	200	150	230	180	110
45.564	44.571	45.5439	44.5568	45.5439	44.559	45.53	44.542
43.603	46.610	43.6082	46.618	43.6082	46.618	43.583	46.605
0.0256	0.0219	0.02576	0.02205	0.02718	0.023	0.02717	0.02349
0.0272	0.0220	0.0345	0.0233	0.041	0.1	0.088	0.089
1.50	E+33	I.85E+33		1.16E+34		3.78E+34	
	0.21 78 4.728 122.5 200 4 6 700 45.564 43.603 0.0256 0.0272 1.50	0.21 $0.26$ $789$ $4.728$ $1.731$ $122.5$ $40$ $200$ $200$ $4$ $4$ $6$ $6$ $700$ $400$ $45.564$ $44.571$ $43.603$ $46.610$ $0.0256$ $0.0219$ $0.0272$ $0.0220$ $1.50E+33$	0.21 $0.26$ $0.17$ $789$ $78$ $4.728$ $1.731$ $4.537$ $122.5$ $40$ $53.33$ $200$ $200$ $100$ $4$ $4$ $3$ $6$ $6$ $6$ $700$ $400$ $400$ $45.564$ $44.571$ $45.5439$ $43.603$ $46.610$ $43.6082$ $0.0256$ $0.0219$ $0.02576$ $0.0272$ $0.0220$ $0.0345$ $1.50E+33$ $1.851$	0.21 $0.26$ $0.17$ $0.22$ $789$ $789$ $4.728$ $1.731$ $4.537$ $1.641$ $122.5$ $40$ $53.33$ $13.33$ $200$ $200$ $100$ $200$ $4$ $4$ $3$ $3$ $6$ $6$ $6$ $700$ $400$ $400$ $200$ $45.564$ $44.571$ $45.5439$ $44.5568$ $43.603$ $46.610$ $43.6082$ $46.618$ $0.0256$ $0.0219$ $0.02576$ $0.02205$ $0.0272$ $0.0220$ $0.0345$ $0.0233$ $1.50E+33$ $1.85E+33$ $1.85E+33$	0.21 $0.26$ $0.17$ $0.22$ $0.5$ $789$ $789$ $15$ $4.728$ $1.731$ $4.537$ $1.641$ $4.537$ $122.5$ $40$ $53.33$ $13.33$ $7.5$ $200$ $200$ $100$ $200$ $100$ $4$ $4$ $3$ $3$ $3$ $6$ $6$ $6$ $6$ $5.5$ $700$ $400$ $400$ $200$ $150$ $45.564$ $44.571$ $45.5439$ $44.5568$ $45.5439$ $43.603$ $46.610$ $43.6082$ $46.618$ $43.6082$ $0.0256$ $0.0219$ $0.02576$ $0.0233$ $0.041$ $1.50E+33$ $1.85E+33$ $1.85E+33$ $1.164$	0.210.260.170.220.50.5 $789$ $789$ $1576$ 4.7281.7314.5371.6414.5371.64122.54053.3313.337.51820020010020010020044333366665.5570040040020015023045.56444.57145.543944.556845.543944.55943.60346.61043.608246.61843.608246.6180.02560.02190.025760.022050.027180.0230.02720.02200.3450.02330.0410.11.50E+331.8E+331.16E+34	0.210.260.170.220.50.50.878978915761576154.7281.7314.5371.6414.5371.644.49122.54053.3313.337.51816.22002001002001002008044333266665.555.570040040020015023018045.56444.57145.543944.556845.543944.55945.5343.60346.61043.608246.61843.608246.61843.5830.02560.02190.025760.022050.027180.0230.027170.02720.02200.3450.02330.0410.10.0881.50E+331.85E+331.16E+343.78

## Phase-3 (Early stage) machine parameters

• A few examples of parameter sets observed in the control room

**Operation parameter** sets [24]

[24] D. Zhou, Simulations of beam-beam effects, Talk presented at the 1st SuperKEKB Beam Dynamics Mini-Workshop, KEK, Jul. 17, 2019.

## **Tolerances of IP aberrations**

## Various IP aberrations: the case of e+ beam

- DY\* (Vertical offset)
- **\*\*** Without beam-beam blowup, the dependence is very normal: Gaussian
- \*\* With beam-beam blowup, the dependence becomes non-Gaussian
- \*\* Luminosity drops faster around DY\*=0 because of additional blowup due to nonzero

### v-angle



## **Tolerances of IP aberrations**

### Various IP aberrations: the case of e+ beam

- DY\* (Vertical offset)
- \*\* Dependence on DY\* is special: widened "M" shape
- \*\* With beam-beam blowup, the behavior can become "Quadri Peaks"









# Phase-3 commissioning

- SAD simulations showed that the tolerances would be severe with lattice [25].

## Phase 3-1 machine parameters

### • A few examples of parameter sets observed in the control room

	2019.06.25		2019.07.01		2019.07.01(op1)		2019.07.01(op2)		2019.07.01(op3)	
	HER	LER	HER	LER	HER	LER	HER	LER	HER	LER
I₅ (A)	0.05	0.03	0.8	0.8	0.8	0.8	0.8	0.8	0.8	0.8
# bunch	78	39	1576		1576		1576		1576	
ε <sub>x</sub> (nm)	4.466	1.64	4.49	1.93	4.49	1.93	4.49	1.93	4.49	1.93
ε <sub>y</sub> (pm)"	16.2	6.05	16.2	6.05	40	6.05	16.2	40	40	40
β <sub>x</sub> (mm)	80	80	80	80	80	80	80	80	80	80
β <sub>y</sub> (mm)	2	2	2	2	2	2	2	2	2	2
σ <sub>z</sub> (mm)	5.05	4.66	5.5	5.2	5.5	5.2	5.5	5.2	5.5	5.2
σ <sub>y</sub> (nm)	180	110	180	110	283	110	180	283	283	283
Vx	45.5345	44.542	45.53	44.542	45.53	44.542	45.53	44.542	45.53	44.542
Vy	43.5835	46.606	43.583	46.605	43.583	46.605	43.583	46.605	43.583	46.605
Vs	0.02717	0.02349	0.02717	0.02349	0.02717	0.02349	0.02717	0.02349	0.02717	0.02349
ξ <sub>y</sub> (Geom.)	0.0073	0.012	0.088	0.089	0.057	0.089	0.034	0.089	0.034	0.057
£(Geom.)	1.95	E+32	3.78	E+34	2.63E+34		2.38E+34		I.99E+34	

\*\*Single beam emittance estimated from XRM data

Operation parameter sets [24]

## **Tolerances of IP aberrations with lattice**

### ► Various IP aberrations: the case of e+ beam

• DPY\* (vert. crossing angle) with parameter set of 2019.07.01

**\*\*** Luminosity becomes very sensitive to DPY\* at very small vertical emittance.

\*\* Luminosity drops faster around DPY\*=0 because of additional blowup due to nonzero v-angle

\*\* With lattice, large amplitude particles (beam-beam tail) pick up more nonlinear forces from the final focus system



## **Tolerances of IP aberrations with lattice**

### ► Various IP aberrations: the case of e+ beam

• DY\* (Vertical offset) with parameter set of 2019.07.01

\*\* Luminosity drops faster around DY\*=0 because of additional blowup due to nonzero v-offset

\*\* With lattice, large amplitude particles (beam-beam tail) pick up more nonlinear forces from the final focus system





- Phase-3 commissioning with crab waist
  - Since 2020, crab waist was introduced and led to luminosity boost. Recently, I started to run beam-beam simulations with crab waist and try to compare with experimental observations.
  - With single-beam  $\epsilon_v$  of 22.5 pm, BBSS simulations predict lum. of ~3.75e34 cm<sup>-2</sup>s<sup>-1</sup> without obvious BBHTI. This is compared to the achieved luminosity of 3.0e34 cm<sup>-2</sup>s<sup>-1</sup> in 2021ab run.
  - In BBSS simulations, the crab waist and the single-beam  $\epsilon_{v}$ were also varied. It was found that both of these parameters are essential in determining the luminosity performance.
  - Weak blowup in  $\epsilon_x$  was observed in the control room, but not well-confirmed.

	2021.	05.14	Commonto
	HER	LER	Comments
I <sub>b</sub> (A)	0.68	0.84	
# bunch		74	
ε <sub>x</sub> (nm)	4.6	4.24	w/ IBS
ε <sub>ν</sub> (pm)	22.5	22.5	Estimated from XRM data
β <sub>x</sub> (mm)	60	80	Calculated from lattice
β <sub>y</sub> (mm)	l.	I.	Calculated from lattice
σ₂ (mm)	6	6	w/ bunch lengthening by impeda
σ <sub>y</sub> (μm)	0.15	0.15	Observed from XRM
Vx	45.52989	44.5247	Measured tune of pilot bunch
v <sub>y</sub>	43.59055	46.57279	Measured tune of pilot bunch
Va	0.02719	0.02212	Calculated from lattice
Crab waist	40%	80%	Lattice design

**Operation parameter** set [26]

[26] D. Zhou, Beam dynamics issues: Comparisons of theories, simulations and experiments, Talk presented at the SuperKEKB 2021ab summary meeting, KEK, Jul. 29, 2021.







- Phase-3 commissioning with crab waist
  - Using beam parameter observed on May. 14, 2021, BBSS simulations were done.
  - Simulations showed that the machine seemed to operate round the BBHTI threshold: The blowup of positron  $\sigma_x^*$  in experimental data occurred around the simulated BBHTI threshold.
  - The observed blowup of  $\sigma_v^*$  of both electron and positron beams were complicated (see 24 hours' history of  $\epsilon_v$ ). BBSS simulations cannot reproduce the trends of  $\sigma_v^*$  blowup.
  - Simulations showed working point (.53,.57) is better: Higher BBHTI threshold and weaker beam-size blowup.

	2021.	05.14	
	HER		Comments
I <sub>b</sub> (A)	0.68	0.84	
# bunch	H	74	
ε <sub>x</sub> (nm)	4.6	4.24	w/ IBS
ε <sub>v</sub> (pm)	22.5	22.5	Estimated from XRM data
β <sub>x</sub> (mm)	60	80	Calculated from lattice
β <sub>y</sub> (mm)	I	I	Calculated from lattice
σ, (mm)	6	6	w/ bunch lengthening by impeda
σ <sub>y</sub> (μm)	0.15	0.15	Observed from XRM
Vx	45.52989	44.5247	Measured tune of pilot bunch
vy	43.59055	46.57279	Measured tune of pilot bunch
V8	0.02719	0.02212	Calculated from lattice
Crab waist	40%	80%	Lattice design

**Operation parameter** set [26]

[26] D. Zhou, Beam dynamics issues: Comparisons of theories, simulations and experiments, Talk presented at the SuperKEKB 2021ab summary meeting, KEK, Jul. 29, 2021.











- Phase-3 commissioning with crab waist
  - On Jul. 1st, 2021, a machine study was done with high bunchcurrents for collision. Strong blowup in LER  $\sigma_v^*$  and obvious blowup in LER  $\sigma_x^*$  were observed in experiment.
  - BBSS simulations were done to compare the experimental observations. With strong BBHTI and assumed bunch lengthening, the simulated slope of specific luminosity seemed to agree with experimental data.
  - Parameters such as  $\epsilon_{y}$ ,  $\sigma_{z}$ , and  $\nu_{x,y}$  were varied in simulations. Turned out that the design working point (.53,.57) has weaker BBHTI, giving high luminosity. BBHTI seemed to play an important role.

	2021.	07.01	
	HER	LER	Comments
l= (A)	le	1.255* <u>le</u>	
# bunch	393		
ε× (nm)	4.6	4.0	w/ IBS
ε <sub>v</sub> (pm)	18	18	Single beam (Estimated from XRM data)
β <b>∝ (mm</b> )	60	80	Calculated from lattice
β <sub>7</sub> (mm)	I	I	Calculated from lattice
σ₂ (mm)	5.05	4.84	Natural bunch length (w/o MWI)
Vx	45.532	44.525	Measured tune of pilot bunch
Vy	43.582	46.593	Measured tune of pilot bunch
Vs	0.0272	0.0221	Calculated from lattice
Crab waist	40%	80%	Lattice design

**Operation parameter** sets [26]



Luminosity history panel seen in SuperKEKB control room

- Phase-3 commissioning with crab waist
  - BBSS simulations showed clear BBHTI threshold. The threshold is sensitive to  $\sigma_z$  and  $\nu_x$ , not sensitive to  $\epsilon_v$  and  $\nu_v$ . This is consistent with BBHTI theory (K. Ohmi et al.)
  - Simulations showed that BBHTI makes vertical emittance growth more severe.
  - From simulations, careful choice of working point can relax BBHTI by increasing its threshold.
  - Experiment phenomena are quite complicated. It was hard to determine the BBHTI threshold. Blowup of  $\sigma_v^*$  was much different from simulations. The two beams had unbalanced blowup.

	2021.	07.01	
	HER LER		Comments
⊫ (A)	le	1.255* <u>le</u>	
# bunch	39	93	
ε× (nm)	4.6	4.0	w/ IBS
ε <sub>y</sub> (pm)	18	18	Single beam (Estimated from XRM o
β <sub>x</sub> (mm)	60	80	Calculated from lattice
β <sub>7</sub> (mm)	I	I	Calculated from lattice
σz (mm)	5.05	4.84	Natural bunch length (w/o MWI
٧x	45.532	44.525	Measured tune of pilot bunch
٧y	43.582	46.593	Measured tune of pilot bunch
Vs	0.0272	0.0221	Calculated from lattice
Crab waist	40%	80%	Lattice design

**Operation parameter** sets [26]



Luminosity history panel seen in SuperKEKB control room





# Summary

- Beam-beam simulations with design lattice  $\bullet$ 
  - Interplay of beam-beam (w/o crab waist) and lattice nonlinearity was found to cause severe luminosity loss.
  - The IR nonlinearity was analyzed and found to be the main source of luminosity loss.
- Beam-beam simulations with Phase-2 and Phase-3 machine parameters  $\bullet$ 
  - Without crab waist, beam-beam resonances set a strong limit in luminosity performance.
  - BBHTI seems to be important, but not confirmed yet. Simulations showed that careful choice of working point can relax BBHTI.
  - Both simulations and experiments showed crab waist is effective in suppressing beam blowup and boosting luminosity.



# References

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