#### 超新星爆発シミュレーションのための 球対称ニュートリノ輻射流体計算のGPUによる高速化 (II)

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### Introduction

#### Core collapsed supernovae

- Large scale numerical simulation is essential to understand explosion mechanism
  - Hydrodynamics
  - Boltzmann equation for neutrino transport
  - General relativity
  - Equation of state of dense matter, neutrino reactions
- Fully 6D simulations are currently restrictive
  - Dimensionality plays an essential role for explosion
  - Approximations are often used for 2D/3D systematics
  - Acceleration of full 2D/3D simulations are waited
- Spherically symmetric system
  - Basis for 2D/3D simulations and observations
  - Systematic survey of massive stars is necessary
  - 1<sup>st</sup> principle calculation (Full GR + Hydro + Boltzmann)



#### Introduction

#### High performance computing

- Two trends of architecture
- Massively parallel supercomputers
  - K-computer  $\rightarrow$  Post-K (2021)
  - Intel Xeon Phi
  - Preparation for the next generation SC is underway
- Arithmetic accelerators
  - GPUs, Pezy-SC
  - CPUs + devices: Heterogeneous architecture
  - Currently not widely used in SN simulations in spite of large potential

This work: for establishing techniques to exploit heterogeneous architectures for supernova simulations

• Application to spherically symmetric system as a testbed



- Numerical setup
  - GR Lagrangian hydrodynamics +  $S_N$  + Implicit scheme

S. Yamada, ApJ 475 (1997) 720, A&A 344 (1999) 533

- Ar every step of time evolution, solve a linear equation system
- BiCGStab algorithm for a block tridiagonal matrix

$$M = \begin{pmatrix} B_1 & C_1 & 0 & \dots & & \\ A_2 & B_2 & C_2 & 0 & & \\ 0 & A_3 & B_3 & C_3 & & \\ \vdots & & \ddots & \ddots & \ddots & 0 & \\ & & 0 & A_{n-1} & B_{n-1} & C_{n-1} & \\ 0 & \dots & 0 & A_n & B_n \end{pmatrix}$$

- Rank of each block matrix:  $N_{max} = N_{E_{\nu}} \cdot N_{ang} \cdot N_{\nu} + N_{hyd}$
- Weighted Jacobi preconditioner

A. Imakura et al. JSIAM Letter 4 (2012) 41

$$x_{k+1} = \omega \left[ -M_D^{-1} (M - M_D) x_k + M_D^{-1} b \right]$$

- This linear equation solver is the first target of offloading

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### Implementation

- Offloading scheme
  - Base code: Fortran, MPI parallelized
  - C-code as intermediate code (for later convenience)
  - Device code is implemented with OpenACC
    - Directive-based framework for offloading
    - Portable, compatible with OpenMP 4.0
  - Multi-GPU: each MPI process handles one GPU device
  - Communication via host process
- Two implementations
  - "Native" code
    - Tuned using OpenACC directives
    - Simple assignement of tasks to threads: each thread compute one row
  - cuBLAS code
    - Well-tuned BLAS library provided by NVIDIA
    - CUDA stream for asynchronous execution (in units of block matrix)



## Machines

- Performance is measured on two machines at KEK
  - All arithmetic operations are in double precision
  - Compiler: PGI + OpenMPI + CUDA environment



	Xeon + Kepler	Power8 + Pascal	
Host processor	Intel Xeon Haswell (6 core) x2	IBM Power8 (10 core) x2	
Peak/core (double)	44.8 GFlops/core	22.9 GFlops/core	
GPU	NVIDIA K40 x2	NVIDIA P100 x4	
#core/device (double)	960	1792	
Peak/device (double)	1430 GFlops	4700 GFlops	
Memory size/device	12	16	
Memory B/W	288 GB/s	720 GB/s	
bus	PCIe Gen2 x16	NVLink (40GB/s x2)	

### Problem setup

#### Linear equation

- At each spatial (radial) point r,
  - Neutrino dof.: energy, angle, species
  - Hydrodynamical dof.: velocty etc. (N<sub>hyd</sub>=11)
  - Block tridiagonal matrix, each block is dense

 $\rightarrow$  matrix rank: N<sub>max</sub> = N<sub>E</sub>\*N<sub>ang</sub>\*N<sub>v</sub> + N<sub>hyd</sub>

- $N_r = 256 \rightarrow$  increased to 512, 1024 (preferable)
- Required memory size =  $N_{max}^{2}*4*N_{r}*8$  Byte

set-1 set-2 set-3 set-4 N<sub>r</sub> 256 256 256 256 14 32 N<sub>F</sub> 16 24 N<sub>ang</sub> 6 8 12 16  $N_{\rm v}$ 4 4 4 4 **N**<sub>max</sub> 347 523 1163 2059 1 GB 2 GB 9.3 GB 32 GB memory

Practical choice of N<sub>max</sub>



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*p*-7



•  $M_D^{-1}$ : inverse of diagonal blocks (no communication)



- cuBLAS code tends to be faster at large  $N_{max}$  region (~x1.5)
- On Pascal, ~140 Gflops/device (3% of peak)

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•  $M - M_D$  : subdiagonal blocks (with communication)



- Communication overhead is small ( $\sim 1/N_{max}^*[N_r/N_{dev}]$ )

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### Acceleration of simulation

- How much is simulation accelerated ?
- Time consuming parts in evolution step
  - Dominant: matrix inversion
  - Subdominant: neutrino reaction (1/20 of matrix inv.)
  - Other parts are (currently) negligible

Result: next page

- GPUs largely accelerate the linear solver
- Now dominant operations are:
  - Preparation of matrix (determination of weight parameter)
  - Neutrino reaction
  - $\rightarrow$  Next targets of offloading



## Acceleration of simulation

- Speed up
  - Elapsed time [sec] is measumred for original and offloaded codes
  - Xeon + Kepler: Nrmax=256, Nang=6, Ne=14

Xeon/Kepler (Set-1)	Np = 8	Np = 1	GPU (cuBLAS)
collision	4.9	40.5	
Matrix total	115.5	911.4	46.2
setup	5.1	40.2	40.8
Inversion (11 iter)	110.4	871.2	5.4

- Power8 + Pascal (x4): Nrmax=256, Nang=8, Ne=16

Power/Pascal (Set-2)	Np = 16	Np = 4	GPU (cuBLAS)
collision	8.6	35.1	
Matrix total	168.0	622.5	45.9
setup	11.8	43.7	45.0
Inversion (7 iter)	156.2	578.8	0.9



# Conclusion/outlook

- Conclusion
  - Acceleration of 1D neutrino-radiation hydrodynamics code by GPUs
  - Efficient development by using OpenACC (+cuBLAS)
    - cuBLAS becomes more efficient for larger block matrices, while OpenACC also works well
  - Large speed up for matrix inversion required in implicit scheme
    - GPU performance is not well exploited, but already enough fast
    - Other parts are to be accelerated urgently
    - Simulations with  $N_r$ =1024,  $N_{ang}$ =12,  $N_E$ =24 becomes practical
    - $\rightarrow$  Application to systematic 1D simulations
- Outlook
  - Application to 2D, 3D code
  - Further optimization and improvement of algorithms
  - Pezy-SC

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